NASA LISA Study Team — talk with us about all things gravitational wave!



lisa







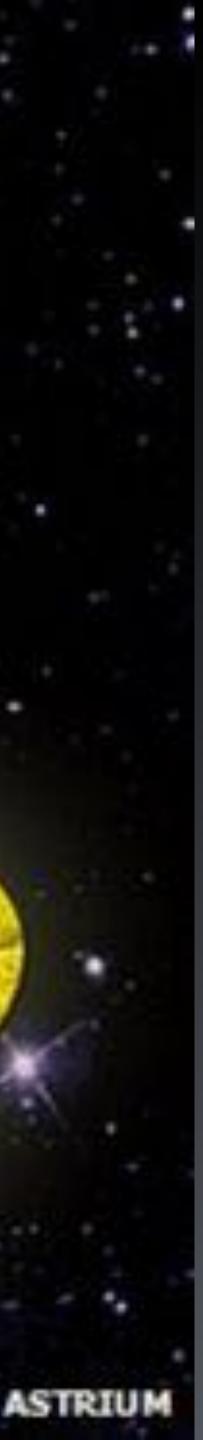








ESA/NASA space-based gravitational wave mission, LISA (proposed launch in 2034)



For objects with orbital periods of seconds — hours... Imagine what you could do with:

> Masses — 0.01 - 1% Distances — 1-10% Spins — 1-10% Spin directions — 10 degrees Eccentricity — 1%







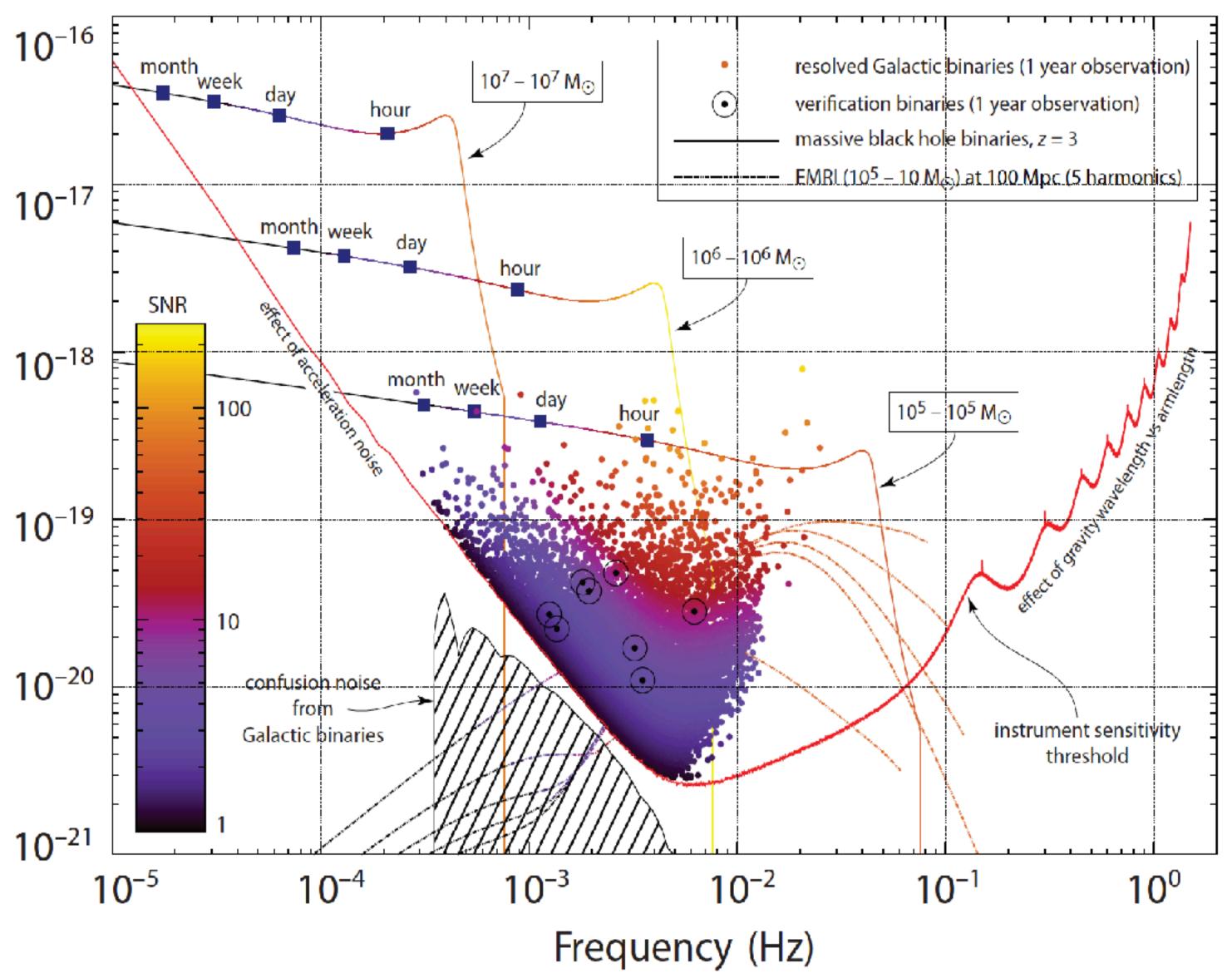
Sky localization — 1 arcmin^2 — 10 deg²

LISA Discovery Space



lisa

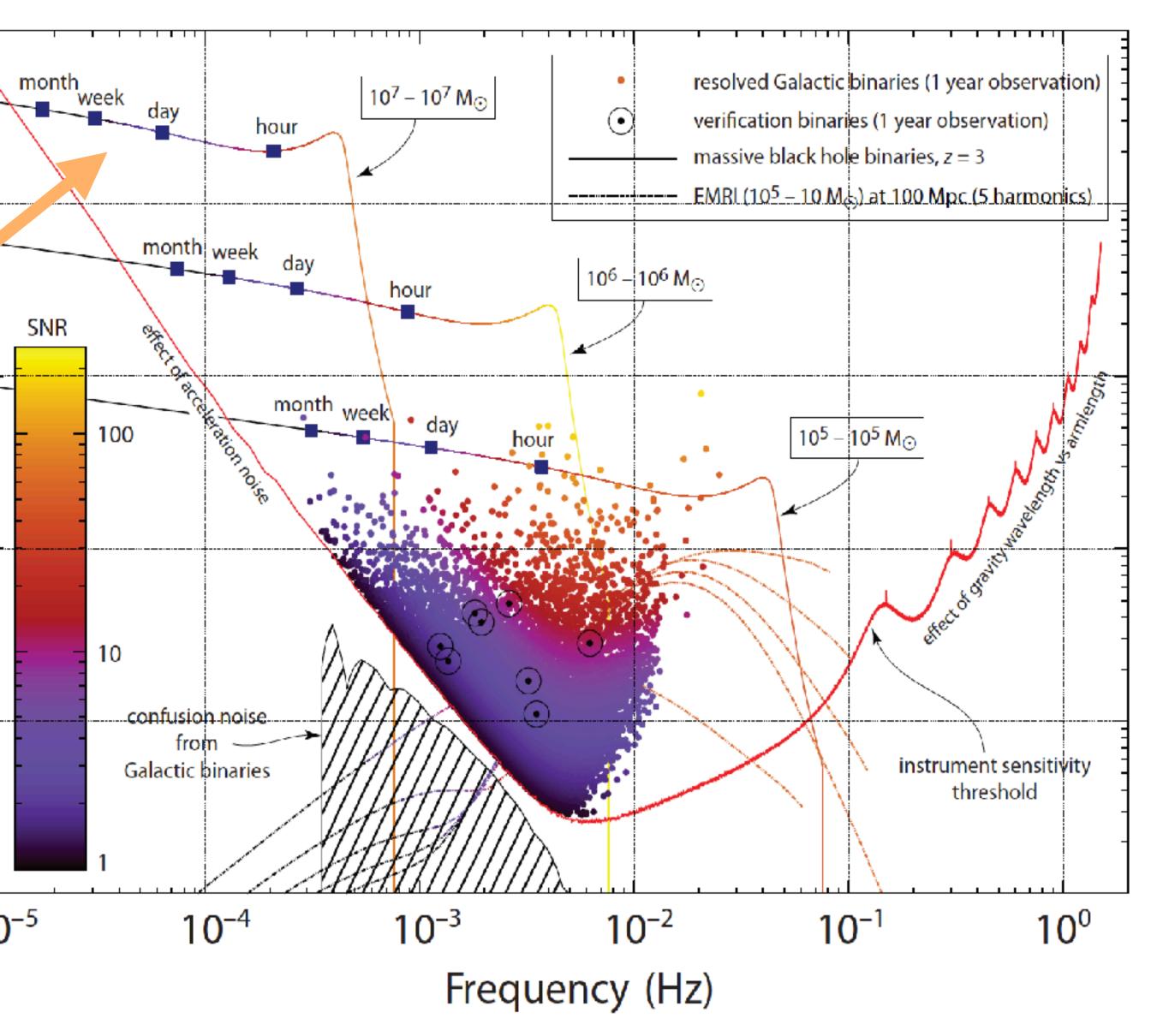
amplitude Characteristic strain

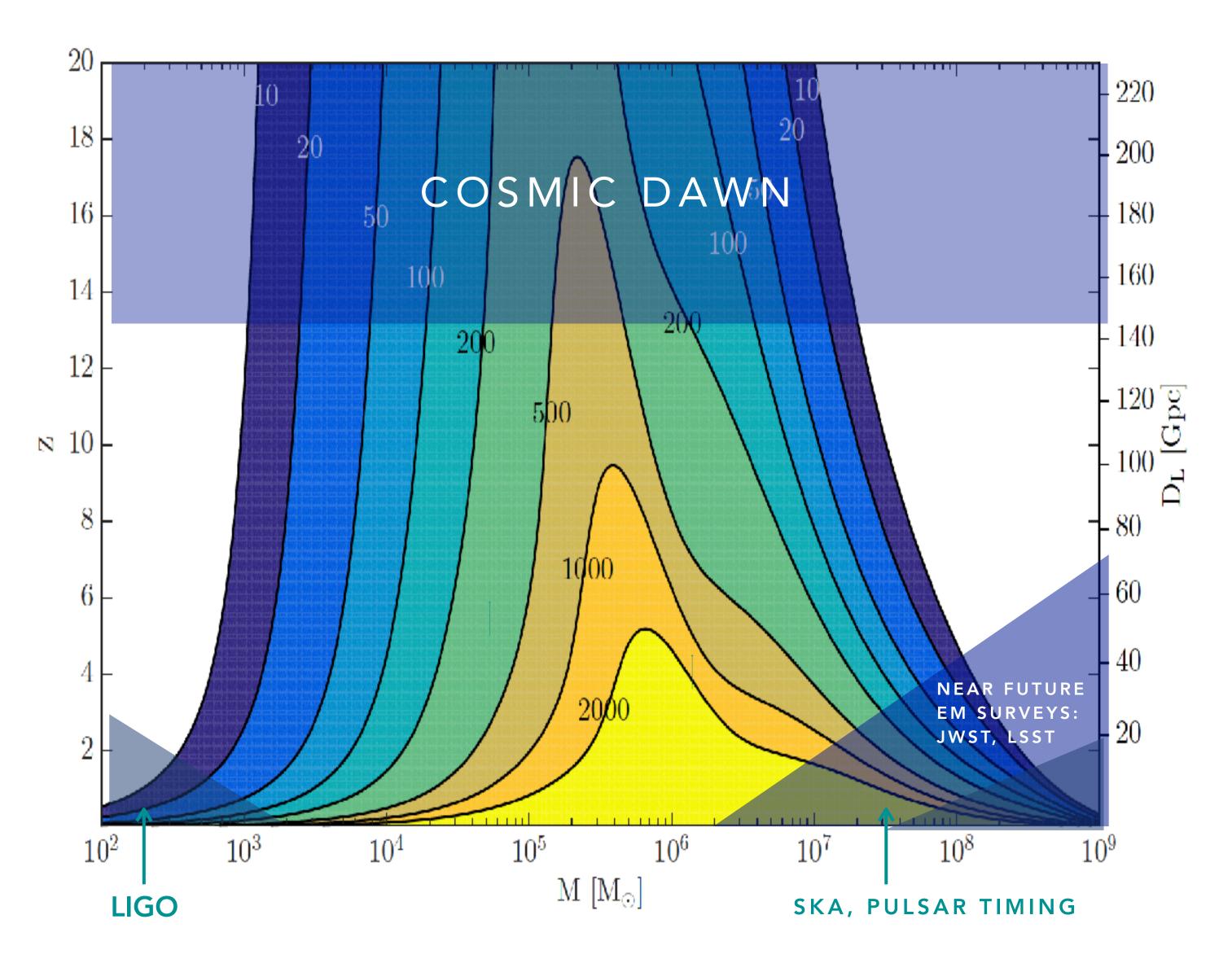


amplitude **10**⁻¹⁷ SNR **10**⁻¹⁸ Mergers of Milky-Way Characteristic strain **10**⁻¹⁹ Class Supermassive 10^{-20} **Black Holes** 10⁻²¹

10⁻¹⁶

10⁻⁵





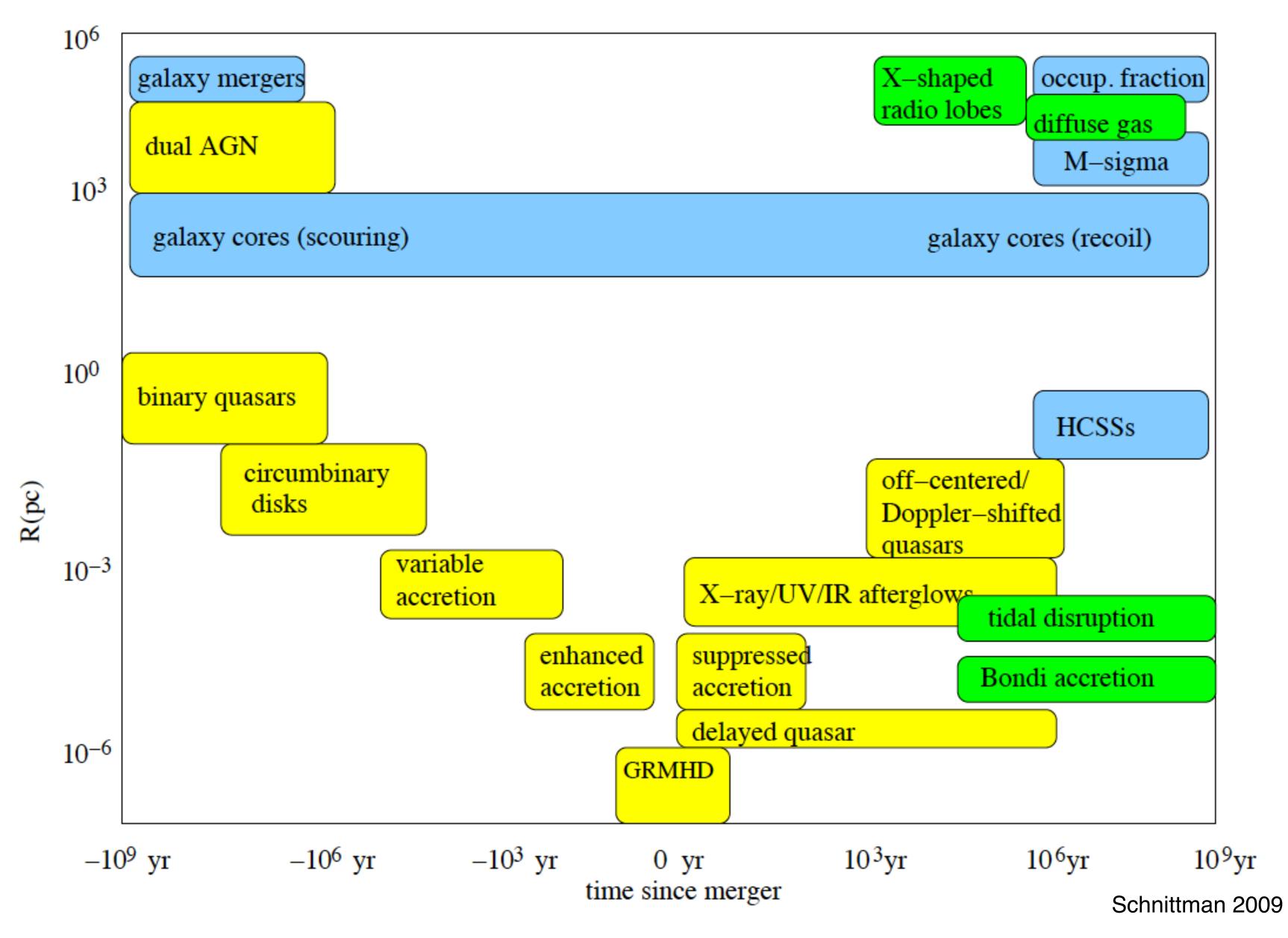
LISA detects the inspiral and merger of intermediate and massive Milky Way-class black holes with huge SNR throughout the observable universe and into the Cosmic Dawn.

What Do We Hope to Measure With LISA and What Do We Hope to Learn?

- \blacklozenge Mergers of binary BHs with M \sim 10⁴ 10⁸ M_ $_{\odot}$
- Trace BH (and galaxy) merger history to z~20 and constrain the origin of BH seeds.
- Complement deep X-ray surveys that trace BH demographics and growth to z~10.
- ♦ Complement UV-O-IR surveys that probe star-formation history and evolution of galaxies to z~8–10 and beyond.
- Identification of individual events for detailed case studies.



While the theoretical and technological dust settles, let's think beyond electromagnetic counterparts



 accurate black hole mass measurements up to z~8 for 10⁵<10⁷ M⊙ - connecting SMBH birth/growth in the reionization era

- the type of galaxy for SMBH hosts

- BH occupation fraction up to z~8 and for Mgal=small

— find evidence of binary black holes (enlist time-domain?)

 look for recoiling AGN (can get 3-d space velocity) — maps to SMBH spin and mass ratio before SMBH merger

mass end)

hypervelocity stars from 3-body scattering out to Coma?

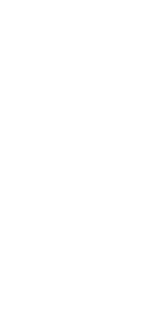
- maybe pulsar planets, nearby highly eccentric and/or hot Jupiter planets (regardless of inclination)

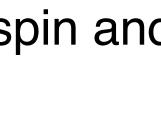
core scouring

LUVOIR can help maximize LISA science, even without electromagnetic counterparts!

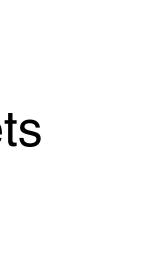
— measure galaxy merger rate to constrain SMBH merger dynamics (esp. @ low

nuclear structure to connect EMRIs to tidal disruption events, and to constrain



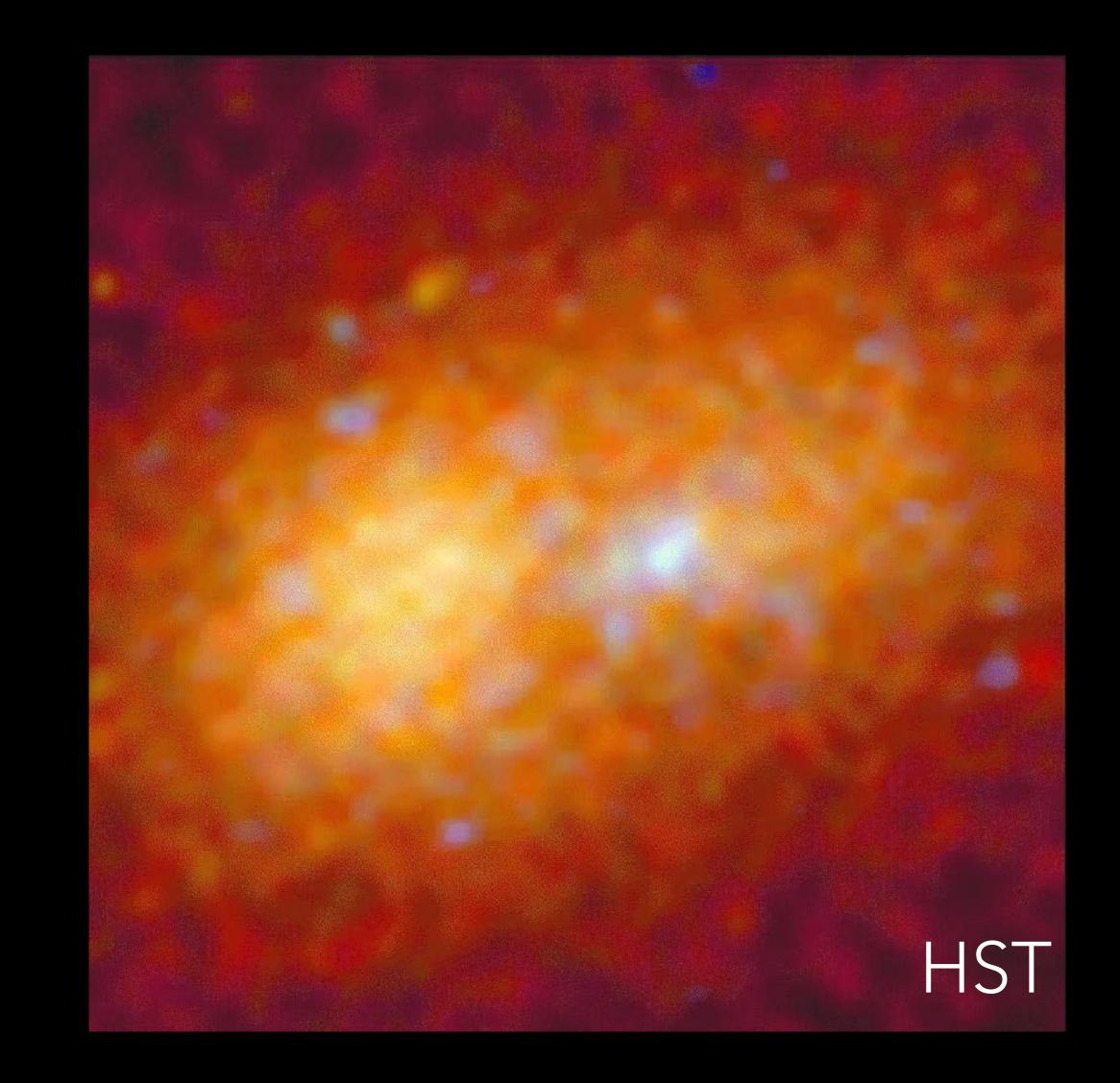


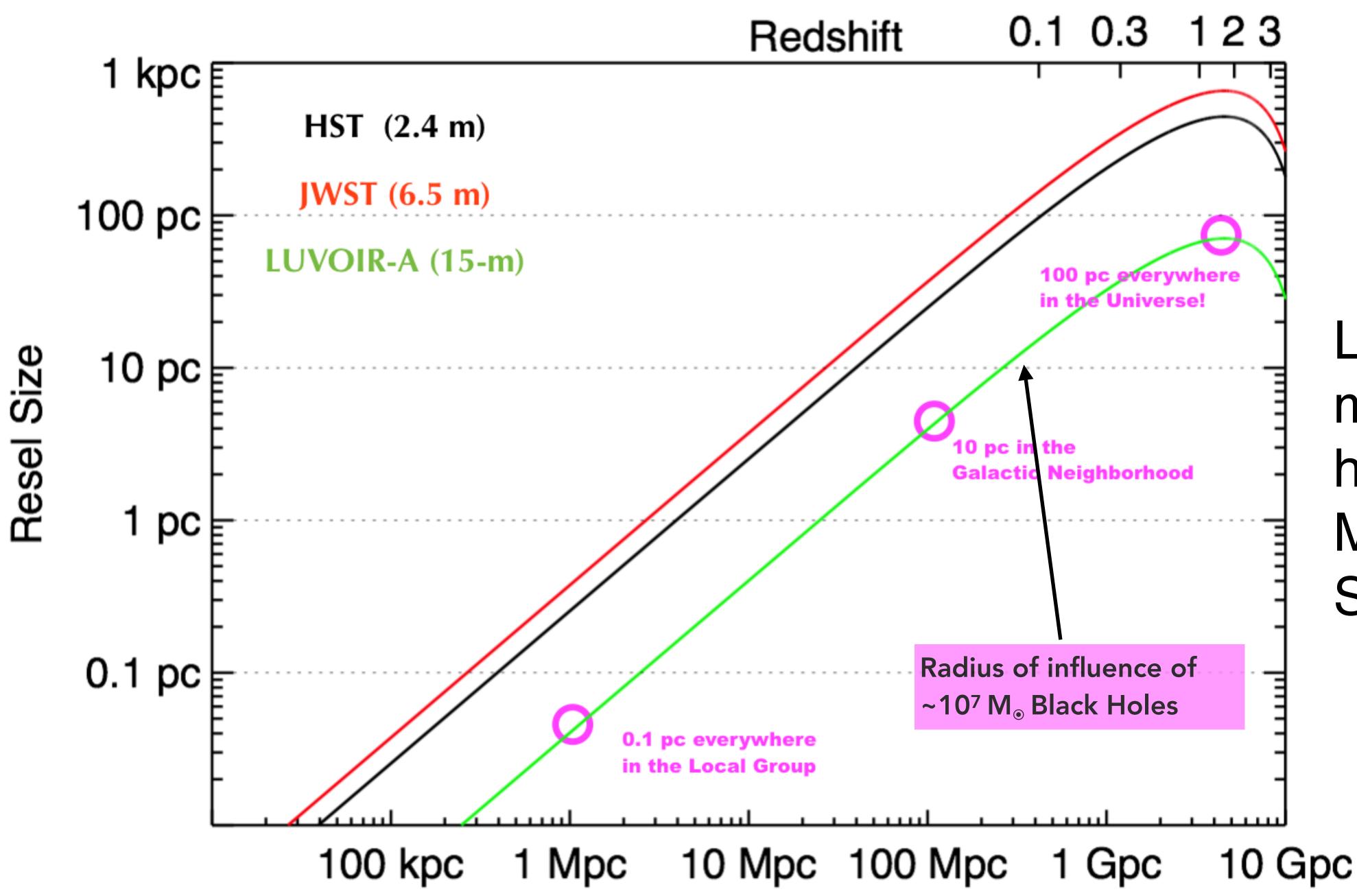






Combine the exquisite resolution of LUVOIR with LISA data

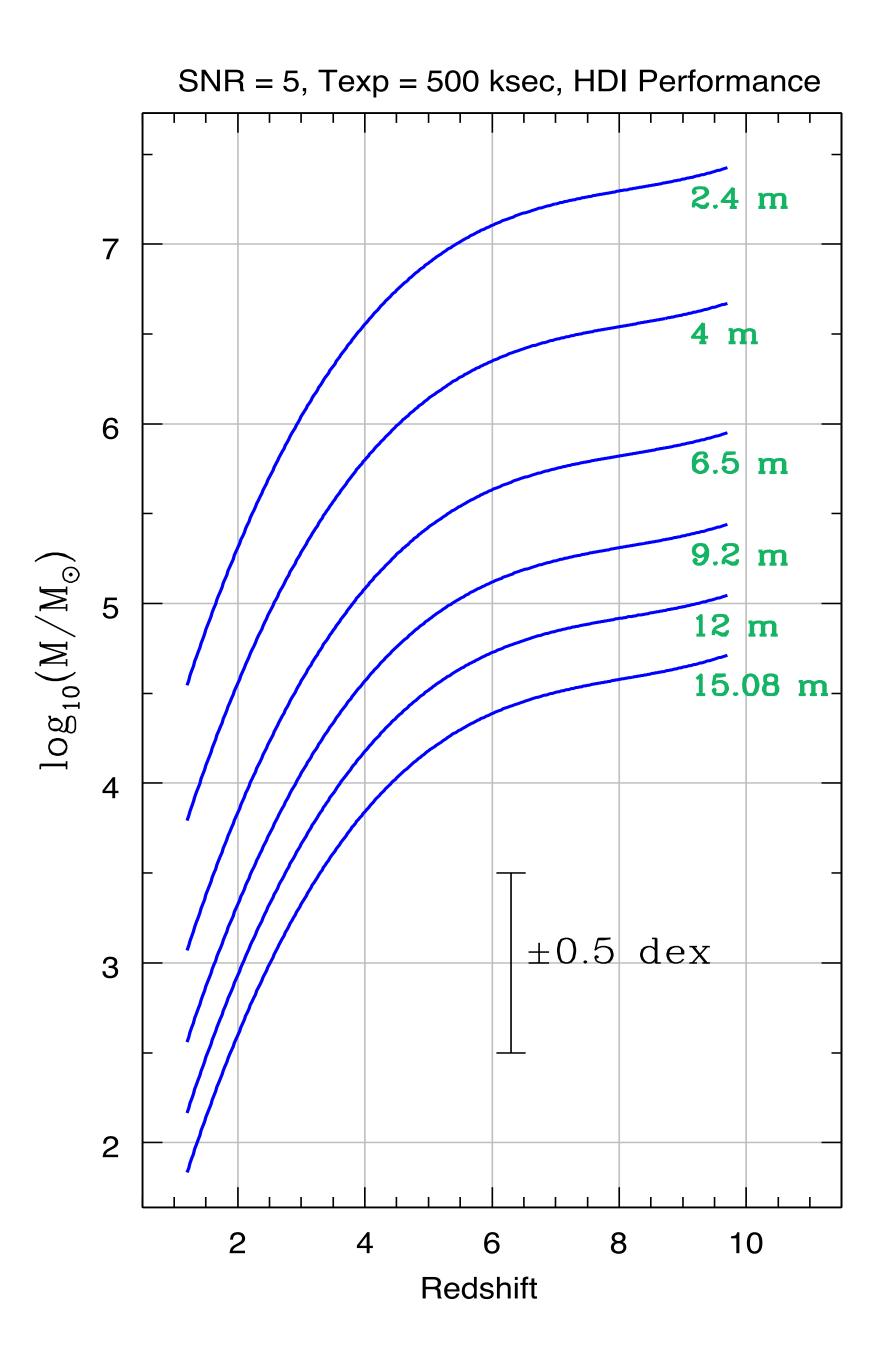


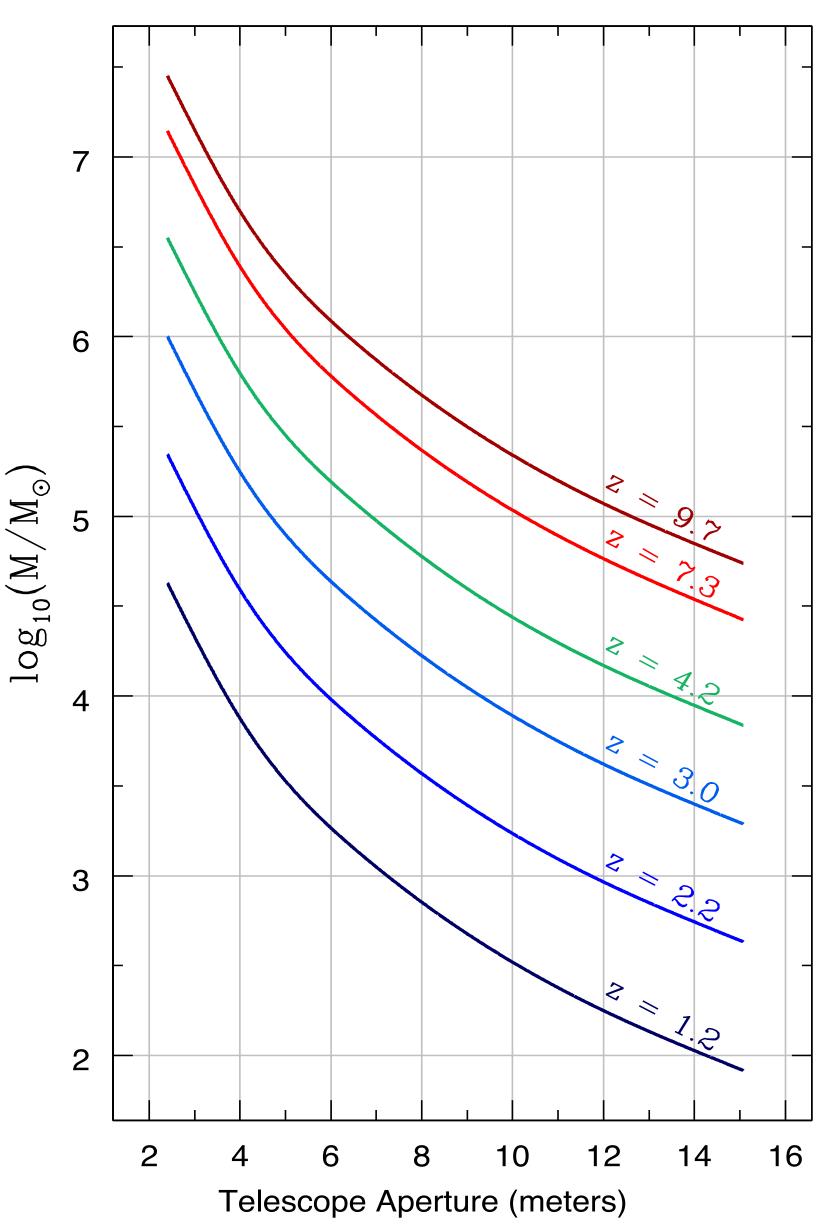


LUVOIR can measure black hole masses for Milky Way SMBHs

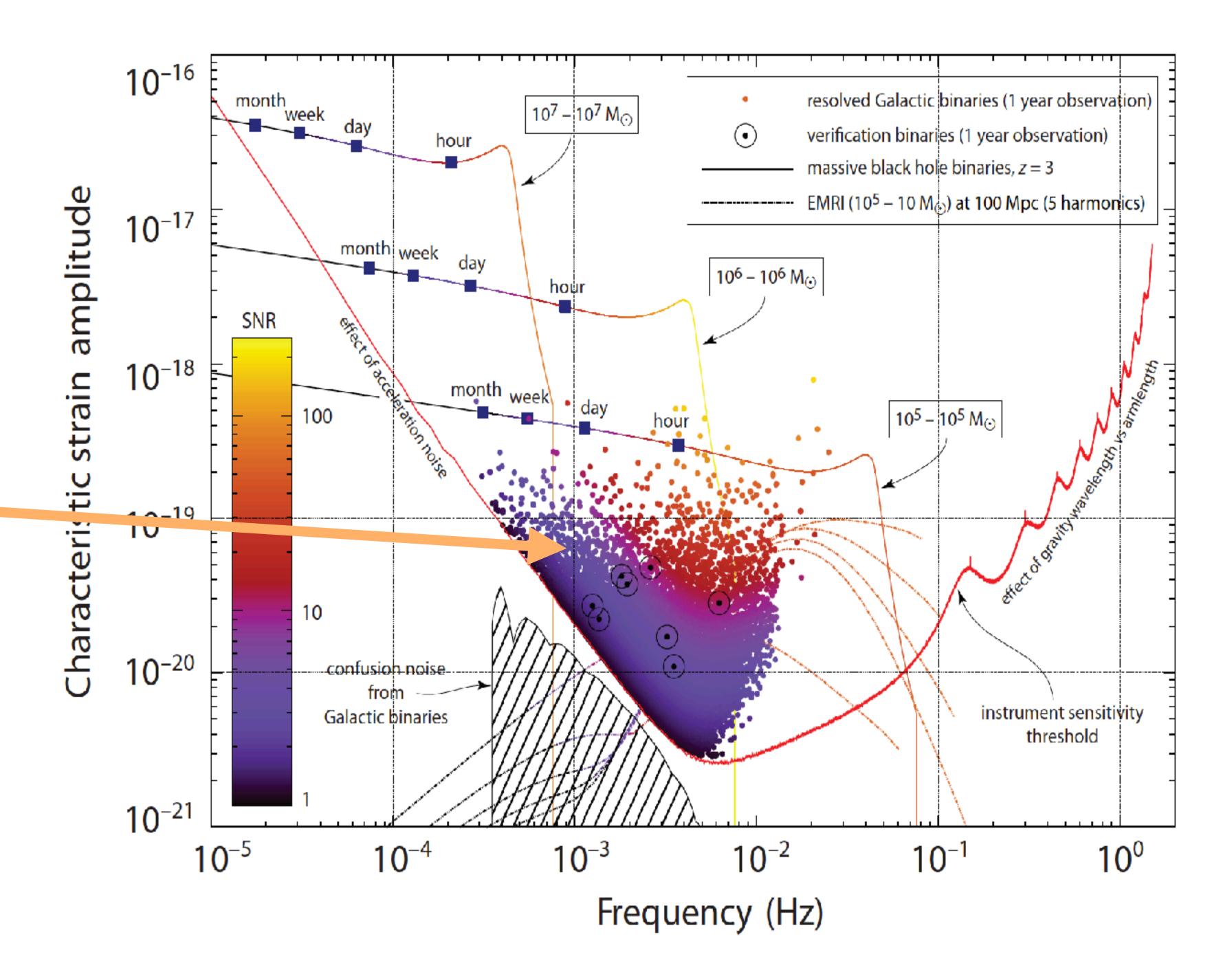


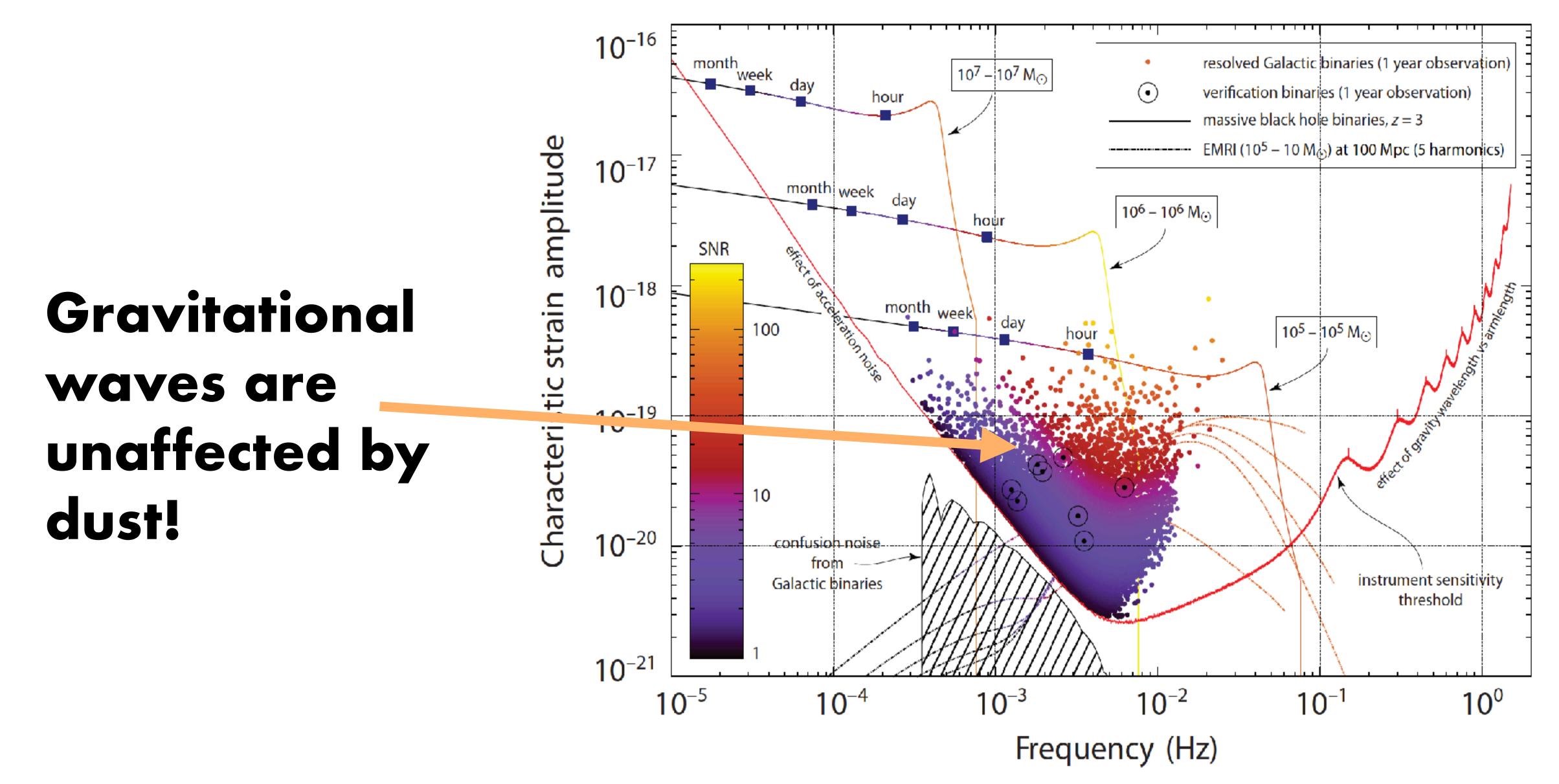
LUVOIR can peer into LISA host galaxies, even for IMBH/seed BH hosts





Millions of close compact object binaries





What do we hope to measure with LISA and what do we hope to learn?

Binary Star Evolution

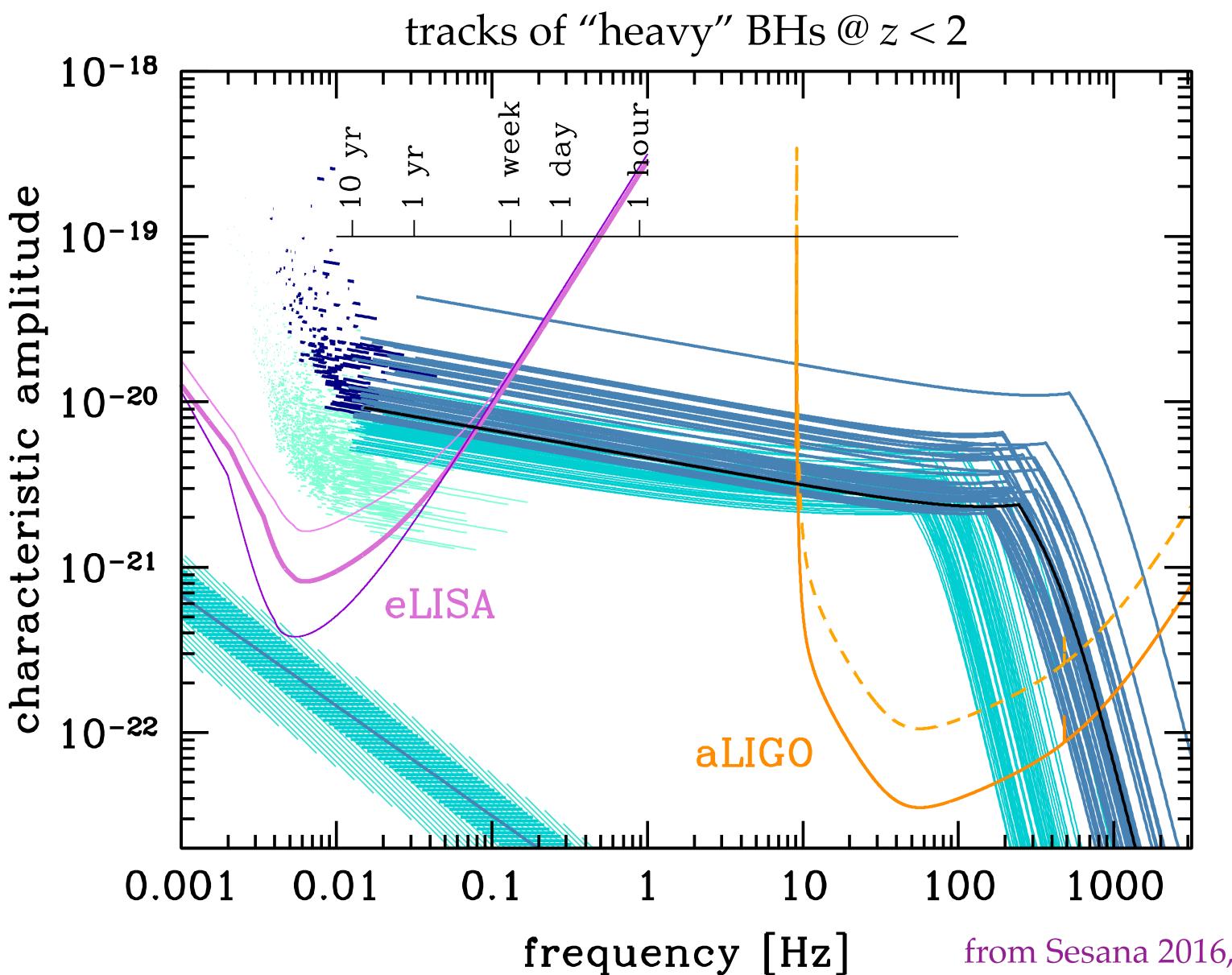
- Census of compact binaries, especially WD+WD
- Determination of binary parameters
- Mapping of old stellar population in Milky Way
- Accretion Physics
 - Obtain system parameters, especially masses and mass transfer rates Combine with known distances/luminosities to constrain accretion flow
 - models

 \rightarrow Preview of LIGO sources (out to $z \sim 2$)



-		
	100	

Extending and complementing LIGO science



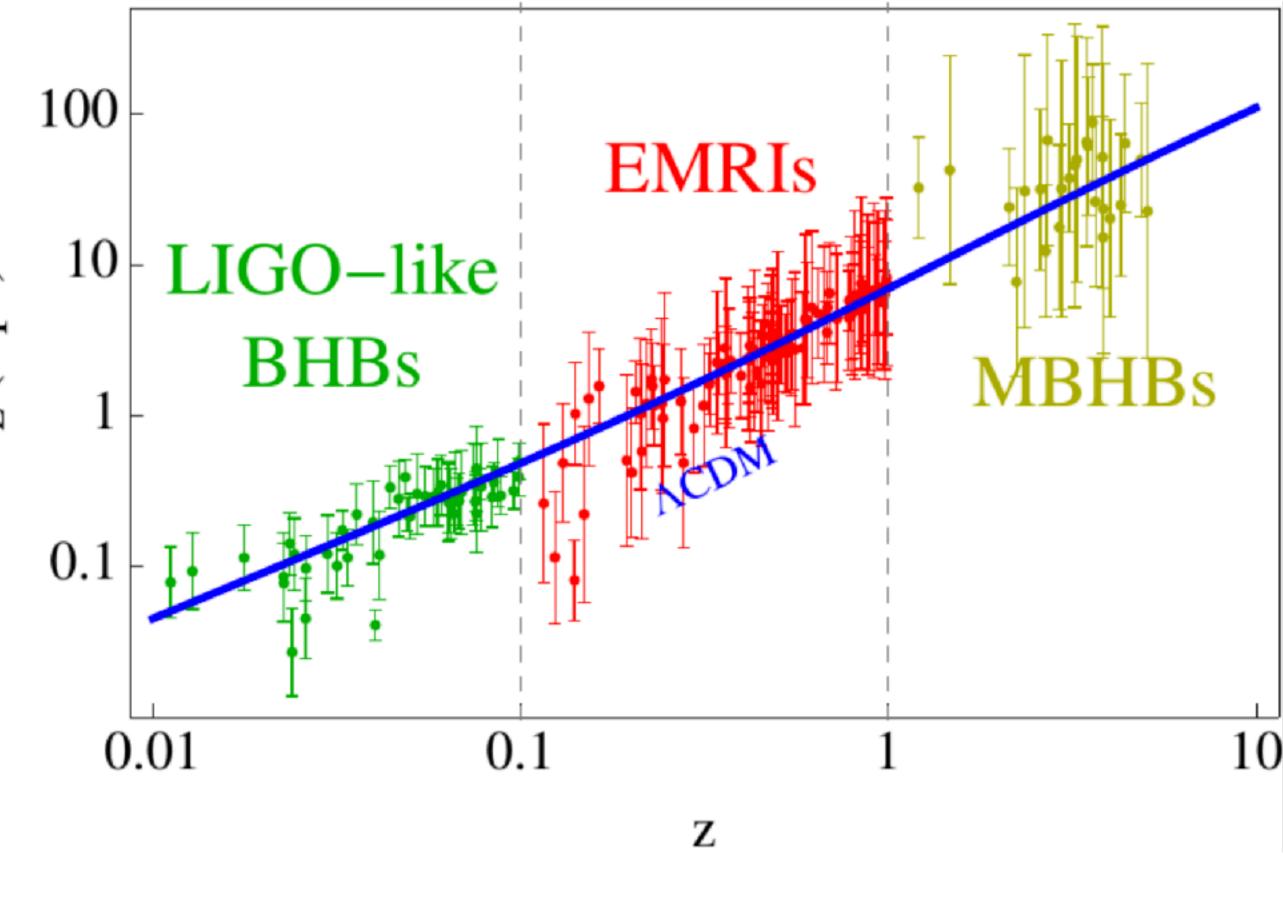
from Sesana 2016, *Phys. Rev. Lett.*, **116**, 231102



Black hole cosmology with standard sirens

- BH mergers as standard sirens
 - chirp rate gives mass
 - mass gives intrinsic amplitude
 - measured amplitude gives distance
- combine with redshift to get H₀
 - Need EM follow-up to identify (or constrain) hosts



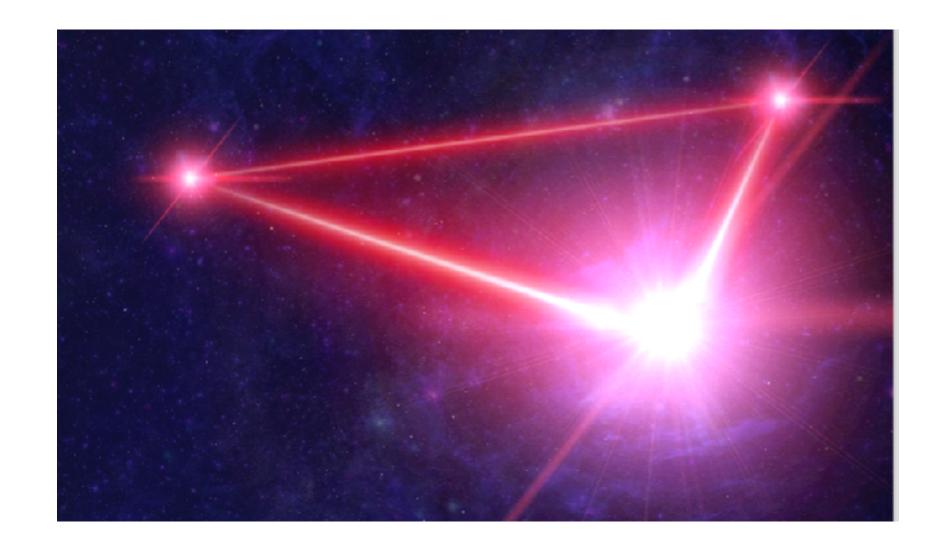


Luminosity distances for simulated catalog of LISA BH binaries (N. Tamanini)



It's a wonderful time to be an astronomer!

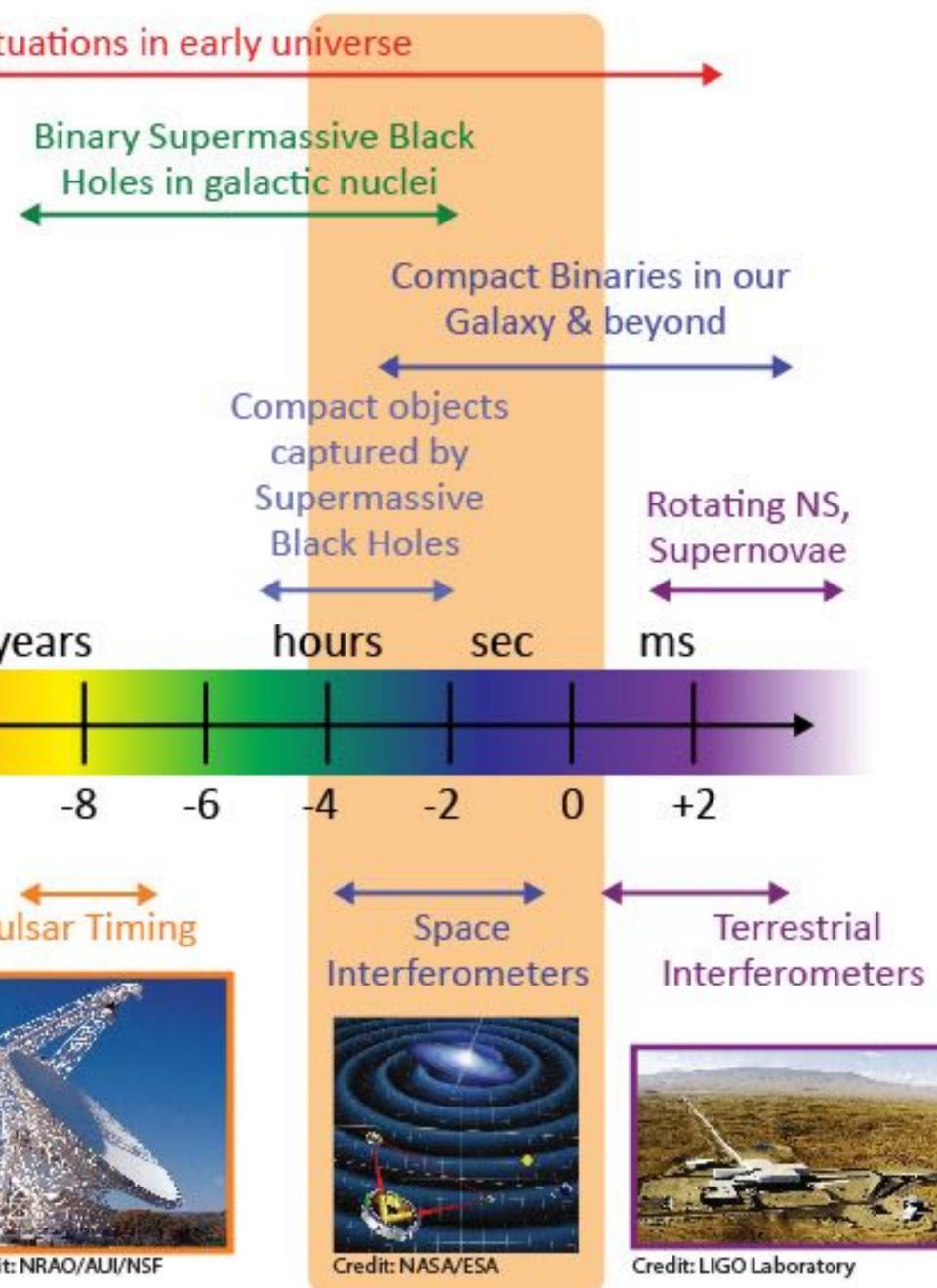
and we need you!



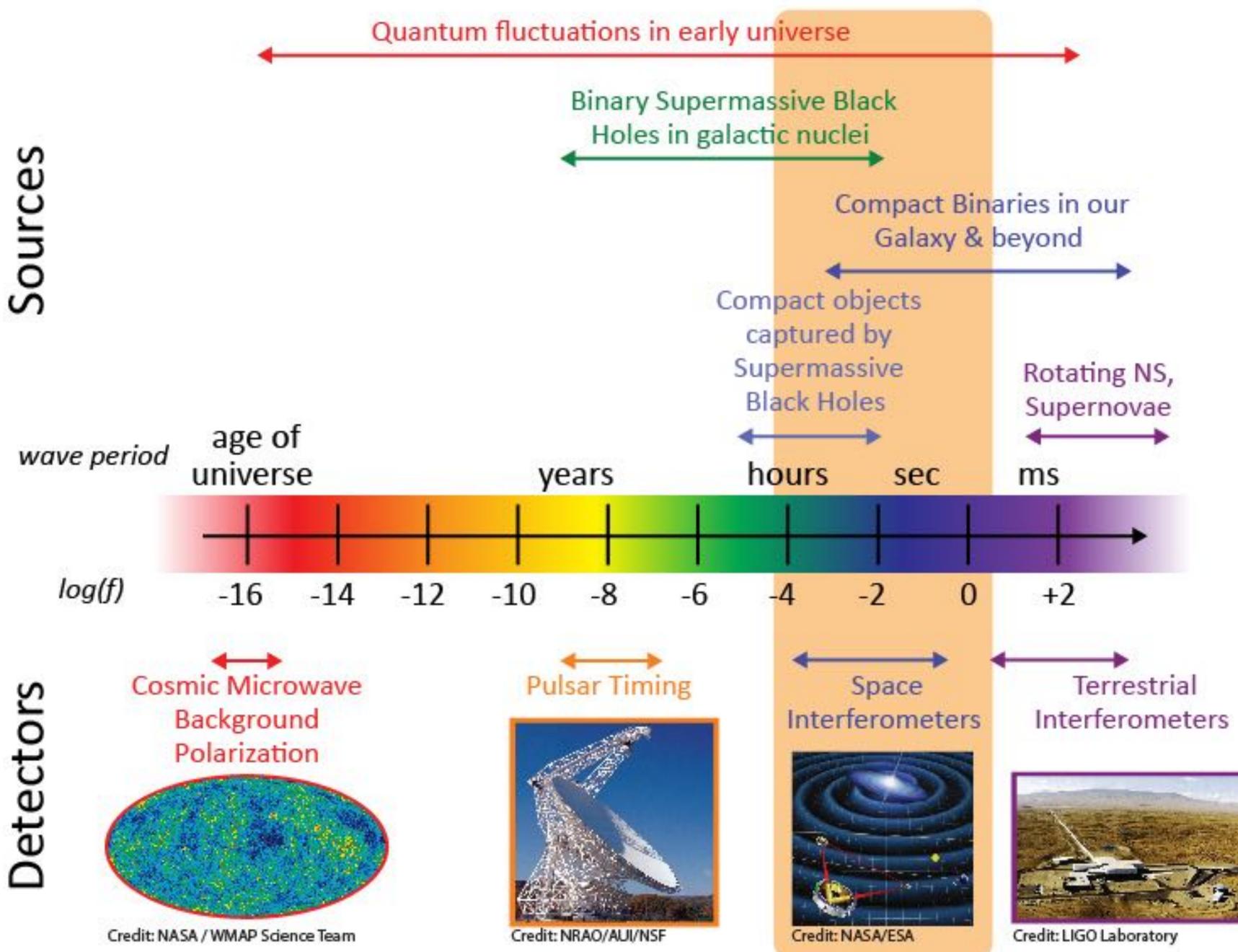
My Fellow Astronomers: LISA is happening! It's time to think about how how to get the most science out of LISA data. We need to build capacity in the brand new field of gravitational wave astronomy

han ks!

The Gravitational Wave Spectrum



Sources



mass ratio before SMBH merger

mass end)

of inclination)

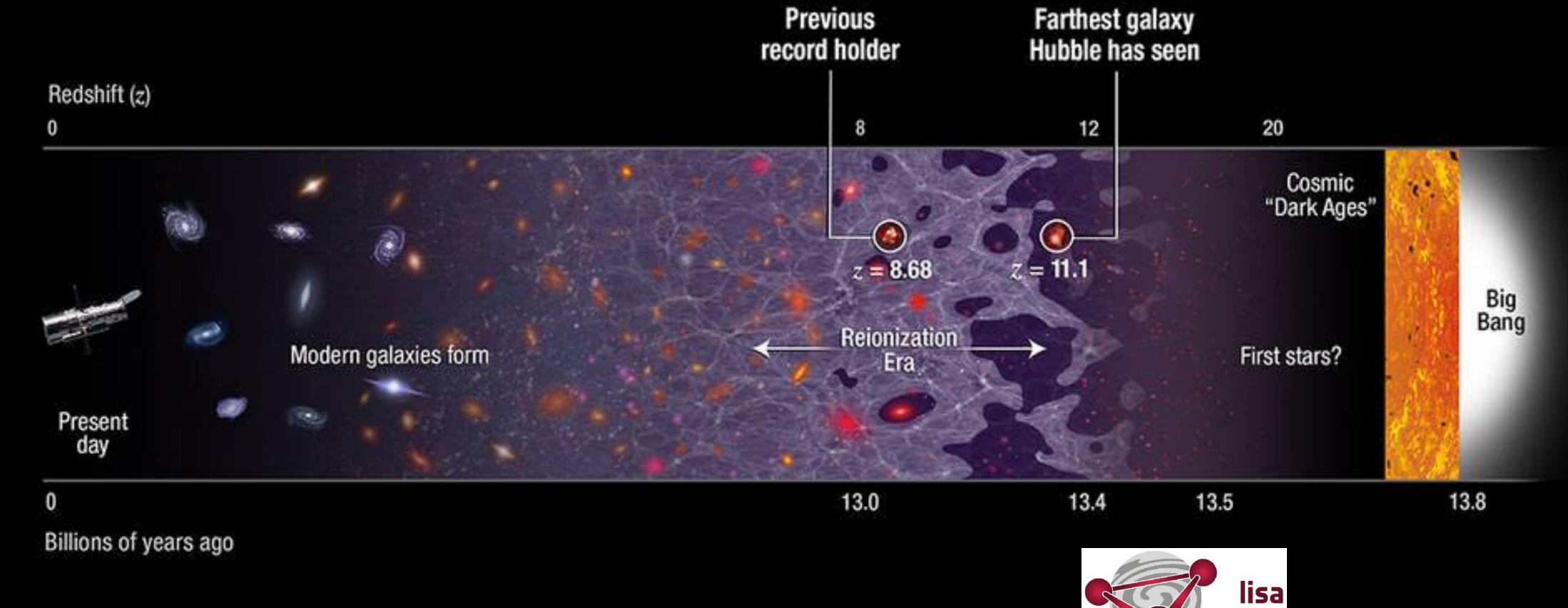
core scouring

LUVOIR can help maximize LISA science, even without electromagnetic counterparts!

- accurate black hole mass measurements up to $z\sim8$ for $10^5 10^7 \, M_{\odot}$
- connecting SMBH birth/growth during the dark ages
- the type of galaxy for SMBH hosts
- BH occupation fraction up to z~8 and for Mgal=small
- find evidence of binary black holes (enlist time-domain?)
- look for recoiling AGN (can get 3-d space velocity) maps to SMBH spin and
- measure galaxy merger rate to constrain SMBH merger dynamics (esp. @ low
- hypervelocity stars from 3-body scattering out to Coma?
- pulsar planets, nearby highly eccentric and/or hot Jupiter planets (regardless)
- nuclear structure to connect EMRIs to tidal disruption events, and to constrain



Hubble spectroscopically confirms farthest galaxy to date



STELLAR REMNANTS ORBITING SMBHS

